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- What are space-time trajectories?
- Why do we need space-time trajectories?
- Space-time trajectory generation
- Space-time trajectories from probabilistic forecast
- Probabilistic and multivariate forecast assessment
- Results
- Conclusions



The most common space-time trajectories are constantly issued by meteorological institutions in the form of ensemble forecasts.

The figure shows temperature timetrajectories at a single grid point generated by the 10 ensemble members of MEPS.



Figure: MetCoOp Ensemble Prediction System, issued 2019-08-15T00:002 for Uppsala, Sweden

### Why do we need space-time trajectories?



Because the spatio-temporal dependence structure provides essential information to operational decision problems [1], for instance:

- Electricity market participation.
- Power system reserve quantification.
- Stochastic model predictive control.
- Probabilistic power flow simulations.

- A regression or machine learning model is often used to map numerical weather prediction forecasts to e.g. PV power or to reduce systematic bias in forecasts.
- A common way to issue probabilistic forecasts is by quantile regression, such that  $q_{\tau} = F^{-1}(\tau)$  is the  $\tau^{\text{th}}$  quantile forecast and  $F^{-1}$  an inverse cumulative distribution function (CDF).
- By choosing e.g.  $\tau = [0.10, 0.30, ..., 0.90]$  it is possible to issue a predictive discrete inverse CDF:



Figure: Example of a predictive discrete inverse CDF.

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• Drawing a random number u from U[0,1] and interpolating, it is possible to generate a value from the discrete inverse CDF:



Figure: Generating a value from the discrete inverse CDF using a random number (0.82) and linear interpolation.

• This is known as the inverse probability integral transform:  $x = F^{-1}(u)$ .

**PVPS** 

• When probabilistic forecasts are issued for z = 1, ..., Z locations and k = 1, ..., K horizons,  $D = Z \times K$  random numbers are required to generate 1 space-time trajectory.

It is important that the uniform random numbers accurately represent the spatiotemporal correlation. Consider the purely temporal example  $(5 \times K)$ :

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**S**dVd



- Consider random variables  $(X_1, ..., X_D)$  with CDFs  $F_{X_1}, ..., F_{X_D}$ .
- Sklar's theorem states that every multivariate CDF can be expressed using the marginals and a copula  $C: F_{X_1,...,X_D}(X_1,...,X_D) = C\left(F_{X_1}(X_1),...,F_{X_D}(X_D)\right)$ . [3]
- The copula can be written as:  $C(u_1, ..., u_D) = F_{X_1, ..., X_D} \left( F_{X_1}^{-1}(u_1), ..., F_{X_D}^{-1}(u_D) \right).$
- Then, it is possible to sample random uniform numbers  $(U_1, ..., U_D)$  and generate space-time trajectories.

### **Data and approach**

• Aim: compare suitable copulas for multivariate solar forecasts.





Figure: Oahu pyranometer network [4]. The red arrow indicates the prevailing wind direction. From [2].



Figure: Flowchart that presents the methodology for selecting the copula (offline) and how it is used in an online setting or on the available test dataset



## Probabilistic and multivariate forecast assessment

- Probabilistic and multivariate forecasts should be calibrated for optimal decision-making processes.
- A flat rank histogram is a necessary condition for calibration, meaning that—on average—it is equiprobable for any ensemble member to predict the observation.
- Since there are *D* dimensions instead of 1, dedicated 'prerank' functions have been proposed that result in the average and band depth rank histograms [4].

$$\begin{pmatrix} y_1 & y_2 & \cdots & y_D \\ x_{1,1} & x_{1,2} & \cdots & x_{1,D} \\ x_{2,1} & x_{2,2} & \cdots & x_{2,D} \\ \vdots & \vdots & \ddots & \vdots \\ x_{S,1} & x_{S,2} & \cdots & x_{S,D} \end{pmatrix} \text{ instead of } \begin{pmatrix} y_1 \\ x_1 \\ x_2 \\ \vdots \\ x_S \end{pmatrix}$$

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**Results** 

- In this case, the Gaussian copula is not flexible enough to model the spatio-temporal dependence structure.
- The Student-*t* copula results in too high correlation or underdispersed trajectories.
- The type of miscalibration from the rank histograms for the Clayton copula is inconclusive except that calibration overall is quite poor.
- The empirical copula produced slightly underdispersed trajectories caused by the probabilistic forecasts.

Figure: The multivariate rank histograms organized by copula (columns) and prerank function. The top row presents the average rank histogram and the bottom row represents the band depth rank histogram [2].







- Space-time trajectories are an important input to decision-making processes where the spatio-temporal relationship contains valuable information.
- Space-time trajectories can be created through several techniques but we focused on probabilistic forecasts and copulas.
- A copula is a versatile tool that allows modeling the dependence structure and marginal distributions separately.
- We found that the parametric copulas (Gaussian, Student-*t* and Clayton) are in this case not flexible enough for the relatively large number of dimensions.
- The empirical copula showed better performance, which is probably because it is nonparametric.

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## Thank you for your attention

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#### **References**



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